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# Direct evidence of causality between chemical purity and band-edge potential fluctuations in nanoparticle ink-based $Cu_2ZnSn(S,Se)_4$ solar cells

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#### Abstract

Kesterite solar cells based on chalcogenide  $Cu_2ZnSn(S,Se)_4$  (CZTSSe) are a viable approach to thin film photovoltaics, utilising Earth-abundant, nontoxic elements. CZTSSe films produced from nanoparticle inks offer a costeffective solution-based method of fabrication. However, improving efficiency in these devices has proved challenging, in part due to the presence of detrimental complex defects within the bulk of the CZTSSe absorber. In this study, the behaviour of nanoparticle-based CZTSSe absorbers and solar cells

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made from relatively low and high quality grade chemicals is investigated with a view to improving cost-effectiveness of the ink-based fabrication process. Photoluminescence (PL) spectroscopy revealed the presence of similar shallow acceptor plus shallow donor states in both low and high purity precursor absorbers. We demonstrate a relationship between the average depth of energy band-edge potential fluctuations and absorber quality where the higher grade chemical precursor-based absorber outperforms the lower purity version. In addition, the low purity precursor absorber had a higher total defect density resulting in a 10 meV increase in the average electrostatic potential fluctuations. Deep level transient spectroscopy (DLTS) in solar devices indicated the presence of detrimental deep defect states in both types of absorber. Notwithstanding the high purity precursor absorber with lower defect density, the power conversion efficiencies of both types of CZTSSe solar cells were similar ( $\sim 5\%$ ), implying an issue other than defects in the absorber bulk inhibits device performance as evidenced by quantum efficiency analysis and current-voltage measurements.

Keywords: Cu<sub>2</sub>ZnSn(S,Se)<sub>4</sub> (CZTSSe), Thin film solar cells, Photoluminescence (PL), Deep Level Transient Spectroscopy (DLTS), Defects

#### 1 1. Introduction

In recent years, the impetus to fabricate low-cost thin-film photovoltaics 2 from naturally abundant, non-toxic elements has focused research inter-3 est on copper-based quaternary chalcogenide kesterites, such as  $Cu_2ZnSnS_4$ 4 (CZTS),  $Cu_2ZnSnSe_4$  (CZTSe) and  $Cu_2ZnSn(S_xSe_{1-x})$  (CZTSSe) (Kumar 5 et al., 2015; Liu et al., 2016). These materials offer a more sustainable al-6 ternative to the current commercially-available thin-film solar cell absorbers 7 CdTe and  $CuIn_yGa_{1-y}Se_2$  (CIGS). They are direct bandgap semiconductors 8 with high absorption coefficients in the visible range ( $\alpha > 10^4 \text{ cm}^{-3}$ ) and 9 CZTSSe has the additional benefit of a tunable bandgap dependent on the 10 S/Se ratio (1.0 eV, x=1 to 1.5 eV, x=0) (He et al., 2012; Siebentritt and 11 Schorr, 2012). Despite a single junction CZTSSe solar cell having a theo-12 retical Shockley-Queisser maximum efficiency limit of 32.2%, polycrystalline 13 CZTSSe devices using a solution approach with reactive hydrazine as the sol-14 vent, have only attained a power conversion efficiency (PCE) of 12.6% (Wang 15 et al., 2013). A less hazardous strategem based on CZTS nanoparticle inks 16 annealed in the presence of Se has achieved device efficiency approaching 17 10% (Hages et al., 2016). Polycrystalline thin-film solar cells are inherently 18 susceptible to the formation of mid-gap defects, which act as electron traps 19 within the bulk of the absorber, and defects at the buffer-absorber inter-20 face which inhibit charge transport. Due to the complexity of the pentanary 21 kesterite crystal structure, a variety of intrinsic lattice defects can form which 22 influence the optical and electronic properties of the CZTSSe absorber (Chen 23 et al., 2013). Density functional theory and first principle calculations have 24 been employed to identify a number of intrinsic defects in bulk CZTSSe, 25

which range from charged point defects such as elemental vacancies, antisites and interstitials to neutral defect complexes. Some defects, like shallow acceptors  $V_{Cu}^{-}$  and  $Cu_{Zn}^{-}$ , are beneficial and are responsible for the *p*-type conductivity of the absorber. Conversely, deleterious defects such as deep donor antisites  $Sn_{Cu}^{3+}$  and  $Sn_{Zn}^{2+}$ , form mid-gap traps which act as effective electron-hole recombination centres.

Band-tails may also be present within the kesterite bulk which have 32 been suggested as contributing to Shockley-Read-Hall (SRH) recombination 33 (Mendis et al., 2013). Two rudimentary processes are understood to cause 34 band-tailing: i) an elevated concentration of highly-compensated charged de-35 fects which result in an electrostatic potential fluctuation of the valence and 36 conduction bands (VB and CB, respectively) and ii) changes in compositional 37 and/or crystalline homogeneity of the absorber, which cause the VB and CB 38 edges to waver, inducing bandgap fluctuations (Gokmen et al., 2013). 39

In this article we present complementary photoluminescence (PL) and 40 deep level transient spectroscopy (DLTS) studies of nanoparticle-based CZTSSe 41 thin films and devices fabricated from low and high purity precursor chem-42 icals (subsequently referred to as film LP and HP, respectively). The evo-43 lution of PL spectra over a range of cryogenic temperatures can be used to 44 elucidate details of shallow defects within both types of CZTSSe absorber, 45 whereas DLTS is used to probe deeper, mid-gap states. Employing both 46 techniques in conjunction provides comprehensive quantitative information 47 on root-mean-square potential fluctuation depths of band egdes  $(\gamma)$ , defect 48 activation energies  $(E_A)$ , defect concentrations  $(N_t)$  and defect capture cross-49 sections  $(\sigma_t)$  of both types of absorber, which is correlated to the structural 50

<sup>51</sup> quality of the material.

#### 52 2. Experimental details

Low and high quality grade metallic sources and solvent oleylamine were 53 used for CZTS nanoparticle synthesis to investigate their influence on the 54 absorbers' behaviour and subsequent solar cells' performance. Low and high 55 purity nanoparticle inks were prepared by varying the metallic sources and 56 solvent as listed in Supp. Table S1. CZTS nanoparticles were fabricated 57 using a hot-injection method where a sulphur-OLA solution was injected 58 into a hot metallic precursors-OLA solution (Qu et al., 2014). The resulting 59 nanoparticle inks were deposited on molybdenum substrates via spin coating 60 to form the CZTS precursor thin film with a thickness around 1  $\mu$ m (Qu 61 et al., 2016a). In order to produce efficient absorbing layers, CZTS precursor 62 thin films were then selenised to introduce grain growth, resulting in CZTSSe 63 absorbers with film thickness  $\sim 1 \ \mu m$ . This process has been described in 64 greater detail in our previous work (Qu et al., 2016b). The crystal structure 65 of both types of film is kesterite in nature determined from X-ray diffraction 66 (XRD). The solar cells were subsequently integrated with a configuration of 67 Mo/CZTSSe/CdS/i-ZnO/ITO/Ni-Al. The CdS buffer layer (~ 55 nm) was 68 deposited using a chemical bath process. *i*-ZnO ( $\sim 35$  nm) and ITO (indium 69 tin oxide,  $\sim 200$  nm) layers were then deposited by magnetron sputtering to 70 act as the transparent oxide layers. Front contact grids which are composed 71 of Ni (~ 50 nm) and Al (~ 1  $\mu$ m) layers were deposited through a shadow 72 mask by electron beam evaporation. Finally, nine  $\sim 0.16 \text{ cm}^2$  cells were de-73 fined by mechanical scribing on each substrate (Campbell et al., 2018). PL 74

spectra were measured using a Horiba Jobin Yvon fully automated spectrom-75 eter fitted with an InGaAs PMT detector cooled to -30°C to reduce noise. 76 A 532 nm continuous wave diode-pumped solid state (CW-DPSS) laser was 77 used as an excitation source. All PL measurements were performed on se-78 lenised absorber layers deposited on Mo coated glass. PL measurements at 79 varying temperatures were performed by placing the sample in a Janis SHI-80 4-2 closed cycle refrigeration cryostat using compressed He gas coupled with 81 a Lakeshore Model 335 temperature controller. Time-resolved photolumines-82 cence (TRPL) experiments were performed at 6 K using 800 nm 1.5 ps laser 83 pulses generated by a tunable mode-locked Ti:Sa laser with a repetition rate 84 of 80 MHz. The laser beam is focused onto the sample on a  $1/e^2$  diameter 85 spot of  $\sim 100 \ \mu m$  and an average power of 20 mW. The PL signal is dispersed 86 by a f-6.5 spectrometer and detected by a synchro-scan Hamamatsu streak 87 camera with an overall time resolution of 15 ps. Deep level transient spec-88 troscopy (DLTS) analysis was performed using a PhysTech FT1230 HERA 89 DLTS system. Emission capacitance transients were recorded using reverse 90 and pulse biases of 5 V and 0 V respectively with a pulse duration of 10 ms. 91 Three transient period widths were recorded 19.2 ms, 192 ms and 480 ms 92 with time constants for the transients being determined via Fourier trans-93 form analysis (DLFTS) (Weiss and Kassing, 1988). Quality of the diodes 94 was assessed prior to DLTS measurements using capacitance-voltage (C-V)95 and current density-voltage (J-V) analysis with C-V measurements being 96 used to calculate shallow acceptor concentration  $N_A$ . A temperature range 97 of 80-300 K was used however samples showed significant variation in reverse 98 capacitance values above 220 K hence only data < 220 K was analyzed. Solar 99

cell J-V parameters were measured using an Abet Technologies solar simu-100 lator at 1-sun (100 mW/cm<sup>2</sup>) illumination equivalent to air mass 1.5 global 101 spectrum with the light power density calibrated using a Si reference cell. 102 External quantum efficiency measurements were performed using a Bentham 103 spectral response system (calibrated using a Si-InGaAs reference cell) under 104 unbiased conditions. C-V parameters were evaluated using an Agilent E4980a 105 LCR meter and a Shimadzu UV-2600 UV-Vis spectrophotometer was used to 106 obtain transmission/reflectiontion data for CZTSSe films on soda-lime glass 107 (SLG). 108

#### 109 3. Results and discussion

CZTSSe is considered a highly-doped and highly-compensated semicon-110 ductor due to a high defect density such that the average distance s between 111 defects is less than the defect Bohr radius (Shklovskij and Efros, 1984). The 112 defects tend to be charged and a random distribution of these unscreened 113 charged defects results in electrostatic potential fluctuations inducing band 114 tails (Bauknecht et al., 2001). The presence of both donor and acceptor 115 defects within the material bulk infers that the material is also highly-116 compensated. Electrostatic potential fluctuations manifest as parallel shifts 117 in valence and conduction band (VB and CB) edges with constant bandgap 118 energy  $E_q$  (Campbell et al., 2018). The consequential band tails with energy 119 lower than  $E_g$  have an exponentially decaying density of states which pro-120 duce a broad asymmetric PL spectrum at room temperature (Levanyuk and 121 Osipov, 1981). Similarly, bandgap fluctuations in the CZTSSe material may 122 be induced by Cu-Zn disorder in the kesterite lattice structure (Schorr, 2011; 123

Choubrac et al., 2013), leading to the formation of non-stoichiometric defect 124 complexes such as  $[V_{Cu} + Zn_{Cu}]$  and  $[Zn_{Sn}^2 + 2Zn_{Cu}]$  (Chen et al., 2013). 125 Slow-cooling at a rate of 10  $^o\mathrm{C/hr}$  after the high-temperature annealing stage 126 of CZTSSe absorber fabrication has been shown to promote clusters of or-127 dered and disordered areas suggesting the random distribution of charged 128 antisite defect  $Cu_{Zn}^{-}$  is reduced (Paris et al., 2014; Schwarz et al., 2013). 129 Consequently, a reduction in local variations of Cu-Zn disorder produces 130 clusters of lower and higher  $E_a$  phases which induces potential fluctuations 131 in the VB and CB of the material. 132

A model describing the defect-related nature of absorption tails caused by electrostatic potential fluctuations has been proposed by Siebentritt et *al.* (Siebentritt et al., 2006). The low energy tail of PL band I(E) is fitted to a Gaussian function,

$$I(E) \sim exp\left(-\frac{(E-E_0)^2}{2\gamma^2}\right) \tag{1}$$

where  $E_0$  is the average emission energy with respect to fluctuating potentials. The magnitude of  $\gamma$  is directly related to the total charged radiative and nonradiative defect density  $N_t$  (Dirnstorfer et al., 1998). Evaluating  $\gamma$  from PL emission spectra therefore allows the absorber total defect density  $N_t$  to be deduced.

Another generally accepted model to describe the behaviour of sub bandgap defects in CZTSSe is the quasi-donor/acceptor-pair (QDAP) model (Yu, 144 1977; Schumacher et al., 2006) which is fundamentally linked to the electrostatic potential model. In this case, the distribution of radiative donor and acceptor states contributing to the potential fluctuation are governed by the Coulombic interaction between the charged defect states. The spectral position of the PL band energy maximum  $E_{PL}$  is described by (Schumacher et al., 2006),

$$E_{PL} = E_g - (E_D + E_A) + \frac{e^2}{4\pi\epsilon_0\epsilon_r s}$$
<sup>(2)</sup>

where  $E_D$  and  $E_A$  are the donor and acceptor energy levels within the CB 150 and VB edges, respectively,  $\epsilon_0$  is permittivity of vacuum,  $\epsilon_r$  is relative per-151 mittivity (or dielectric constant) and s is the separation distance between 152 donor and acceptor. The QDAP process is subject to caveats: (i) the spa-153 tial separation between charged donor and acceptor defects is large enough 154 that the Coulombic interaction between them is negligible and, (ii) all QDAP 155 states are fully occupied when the QDAP PL peak no longer blue-shifts with 156 increasing excitation intensity, indicating maximum Coulombic attraction 157 between defects (Gunawan et al., 2012). At the point of maximum Coulom-158 bic attraction (fully occupied QDAP defects), the blue-shift magnitude of 159 QDAP PL peak  $\Delta E$  is equivalent to the Coulomb potential energy at that 160 point, 161

$$\Delta E = \frac{e^2}{4\pi\epsilon_0\epsilon_r s} \tag{3}$$

Thus, the average separation distance s between acceptor/donor can be derived from Eq.(3) and QDAP defect density  $N_D$  estimated by assuming a uniform distribution of defects within a spherical volume of radius r,

$$r = s = \left(\frac{4\pi N_D}{3}\right)^{-\frac{1}{3}} \tag{4}$$

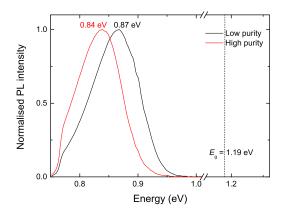


Figure 1: PL of sample LP and HP at 6 K with laser power below PL peak saturation intensity to avoid band-related recombination. The dashed line indicates room temperature bandgap energy  $E_g$  of both types of sample.

The normalised 6 K PL spectra of LP and HP films at excitation intensity 165 just below  $E_{PL}$  peak saturation are illustrated in Fig. 1. The broad asym-166 metric shape of the PL bands for both absorbers show a gradual rise in PL 167 intensity on the low energy side and sharper decline on the high energy side 168 of the PL peak, indicating significant band-tailing in the materials. The PL 169 peak maxima are located at 0.86 and 0.84 eV for LP and HP films respec-170 tively. The energy peaks are considerably red-shifted from the corresponding 171 room temperature bandgap energy of 1.19 eV, determined from EQE data 172 (discussed later in this section) for both types of absorber. Such a large 173 red-shift could be explained by the presence of deep donor defect  $Sn_{Zn}$  with 174 an energy level  $\sim 420$  meV below conduction band minima (CBM) or deep 175 acceptor states  $V_{Sn}$  or  $Cu_{Sn}$  with energy levels ~400 and 420 meV above 176 valence band maxima (VBM), respectively (Chen et al., 2013). The full 177

width at half maximum (FWHM) of the PL spectra exhibit a slight decrease 178 from 96 meV for film LP to 92 meV for film HP, with maximum intensity of 179 the PL band of film HP half the intensity of film LP. The low temperature 180 excitation-dependent intensity of the PL spectra for both absorbers is shown 181 in Fig. 2a. The PL band intensity saturates at lower laser power for film HP 182 and a high energy shoulder at  $\sim 0.92$  eV emerges upon increasing excitation 183 beyond the threshold intensity (see Supp. Fig. S1). Above the threshold, 184 the lower energy peak red-shifts as the high energy shoulder increases in in-185 tensity. Gershon et al. observe similar behaviour in CZTS films and state 186 the high energy shoulder is due to a recombination process associated with 187 extended band states and only appears after all localised (QDAP) states are 188 fully saturated (Gershon et al., 2013). This is further evidence QDAP ra-189 diative recombination is dominant in the CZTSSe films studied here. The 190 appearance of a high energy shoulder was not observed in the PL bands for 191 film LP, indicating all QDAP states are not occupied. An emperical asym-192 metric double sigmoidal function was used to fit the PL spectra in order 193 to evaluate the peak maxima at each excitation intensity (Krustok et al., 194 1999; Yakushev et al., 2017). Both films exhibit a substantial blue-shift of 195 PL peak maxima with increasing laser power until reaching saturation point 196 of PL emission. Fig. 2b shows the evolution of PL band peaks  $E_{PL}$  for 197 the CZTSSe samples over a laser intensity range up to the threshold value. 198 Changes in excitation intensity produce a shift in  $E_{PL}$  ( $\beta$ ) at a rate of  $\beta$ 199 equals 14 and 15 meV/decade and an energy blue-shift magnitude  $\Delta E$  of 30 200 and 37 meV for HP and LP films, respectively. Oscillations in the PL bands 201 around 0.9 eV are associated with water vapour absorption of light. 202

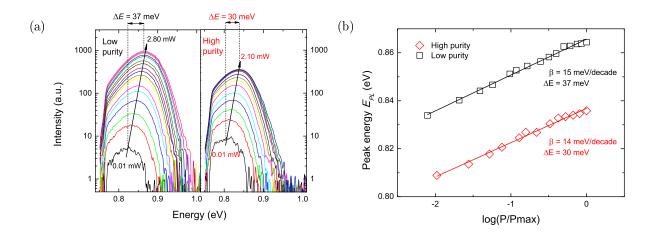


Figure 2: (a) Excitation-dependent PL spectra of sample LP and HP with laser power P up to saturation point of PL emissions together with respective  $\Delta E$  blue-shift values in  $E_PL$  and (b) evolution of PL band maxima with increasing P of LP and HP films with associated shift rates  $\beta$  at 6 K.

Evidence of a significant red-shift in  $E_{PL}$  compared to  $E_g$  together with 203 large  $\Delta E$  plus  $\beta$  values for both LP and HP films indicate QDAP defects 204 are predominantly responsible for potential fluctuations in the valence band 205 maxima and conduction band minima (VBM and CBM) of the CZTSSe ab-206 sorbers. As both films were slow-cooled at a rate of  ${\sim}7^o{\rm C}/{\rm min}$  after anneal-207 ing, it is expected that any bandgap variations contributing to the potential 208 fluctuations should be reduced as the rapid cooling process promotes a more 209 disordered kesterite structure. It is evident that electrostatic potential fluc-210 tuations exist in both films with average amplitude  $\gamma$  values of 48.6 and 58.7 211 meV for films HP and LP respectively which are similar to those previously 212 reported (Gokmen et al., 2013). The QDAP defect density  $N_D$  in film LP 213  $(2.1 \times 10^{18} \text{ cm}^{-3})$  is double that of film HP  $(1.1 \times 10^{18} \text{ cm}^{-3})$  and correlates 214 to a reduction in average defect separation s from 6.0 to 4.8 nm in films HP 215

and LP respectively. Here, electrons and holes are spatially separated and 216 localised in potential wells within the energy band edges and any radiative 217 recombination requires tunneling of charge carriers from one potential well 218 to another. Increasing the separation distance of the defect centres reduces 219 the probability of radiative emission. The concentration of net (free) charge 220 carriers reduces as the charged QDAP defect density increases. Reducing 221 the density of free carriers lowers the screening effect on the charged defects 222 which, in turn, increases perturbation in the depth of the fluctuating poten-223 tial  $\gamma$ . The QDAP density is a measure of the shallow donor and acceptor 224 defect concentration responsible for radiative recombination within the ma-225 terial and may not represent the total defect density  $N_t$  as the presence of 226 deep and/or non-radiative defects is not accounted for. Therefore the total 227 defect density can be estimated from  $\gamma$  values derived from PL spectra of 228 both films (Dirnstorfer et al., 1998). The optical parameters of LP and HP 229 films are outlined in Table 1. 230

Temperature dependence of the PL bands was measured at an excita-231 tion power slightly lower than the threshold value (at which the high energy 232 shoulder emerges) in order to prevent contributions to PL emissions from 233 band-related transitions. PL spectra of both films over a range of cryogenic 234 temperatures are presented in Fig. 3a. It is apparent that PL emissions are 235 quenched at a lower temperature in film HP and both films exhibit a red-shift 236 in  $E_{PL}$  with increasing temperature. The magnitude of the red-shift in  $E_{PL}$ 237  $(\kappa)$  for both types of absorber are shown in Fig. 3b. Initially  $E_{PL}$  red-shifts 238 with rising temperature at a similar rate of 0.05 and 0.09 meV/K for films 239 LP and HP respectively. PL emission from film HP is then quenched at a 240

temperature of ~100 K, whereas  $\kappa$  increases rapidly above ~80 K at a rate 241 of 0.84 meV/K in film LP before emissions are finally quenched at  $\sim$ 140 K. 242 Thermal quenching can be surmised to originate from the thermal depopu-243 lation of defect states and/or the activation of non-radiative recombination 244 centres. Similar behaviour was observed by Grossberg et al. in an optical 245 study of defect clusters in CZTS polycrystals (Grossberg et al., 2014). They 246 state the observed evolution of  $E_{PL}$  with increasing temperature in CZTS 247 polycrystals followed the temperature dependence of the bandgap energy of 248 the material (Sarswat and Free, 2012). Another study by Yakushev et al. 249 revealed comparable temperature trends in PL spectra of CZTSe thin films 250 with varying Cu content (Yakushev et al., 2017). 251

	Low purity	High purity	Defect (Chen et al., 2013)
$\gamma ~({ m meV})$	58.7	48.6	
$\Delta E \ ({\rm meV})$	36.8	29.6	
Defect spacing $s$ (nm)	4.8	6.0	
QDAP density $N_D \ (\text{cm}^{-3})$	$2.1 \ge 10^{18}$	$1.1 \ge 10^{18}$	
Defect density $N_t \ (\mathrm{cm}^{-3})$	$2.3 \ge 10^{19}$	$1.4 \ge 10^{19}$	
$E_a \ ({ m meV})$	$(1) 38.3 \pm 3.8$	$(1)40.1\pm 6.4$	$V_{Cu}$
	$(2) 3.2 \pm 0.5$	$(2) 5.9 \pm 0.5$	$Sn_{Cu}$

Table 1: List of optical parameters of low and high purity CZTSSe samples.

The Varshni equation was used to fit the asymptotic low temperature

252

<sup>253</sup> behaviour of  $E_{PL}$  in both LP and HP films (Varshni, 1967),

$$E_g(T) = E_g(0) - \left(\frac{AT^2}{T+B}\right)$$
(5)

where  $E_g(0)$  is the bandgap of the semiconductor at 0 K and A, B are fitting 254 parameters specific to the absorber material. Using the fitting parameters 255 extracted from application of the Varshni model,  $E_q(0)$  values of 1.24 eV for 256 both films were extrapolated from the room temperature  $E_g$  estimate of 1.19 257 eV.  $E_{PL}$  in both films appear to follow the  $E_g$  trend at lower temperatures 258 until  $\sim 80$  K where it deviates markedly in film LP. If electrostatic potential 259 fluctuations of the VBM and CBM are present in the absorbers due to clusters 260 of charged QDAP defects, shallow and deep potential energy wells which 261 follow the fluctuating band edges are formed. Due to a lack of thermal 262 energy at low temperature, carriers are trapped in shallow wells resulting 263 in incomplete occupation of deeper/least energetic wells. As temperature 264 increases the carriers are energised and liberated to fill the deepest wells with 265 a consequent red-shift in  $E_{PL}$ . PL emission is quenched in both films due 266 to increasing activation of non-radiative mid-gap recombination centres with 267 rising temperature. The large red-shift in  $E_{PL}$  of film LP at temperatures 268 >80 K suggests a different defect is involved in radiative transitions. 269

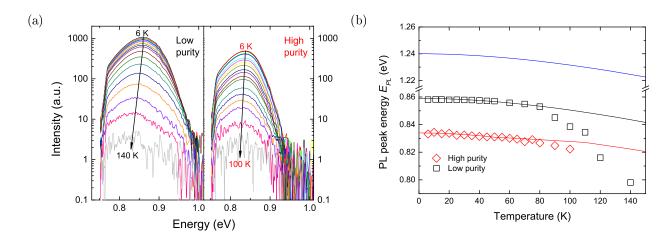


Figure 3: (a) Temperature-dependent PL of sample LP and HP with laser power below PL peak saturation intensity to avoid band-related recombination and (b) evolution of PL band maxima  $E_{PL}$  of LP and HP films with increasing temperature. Eq. (5) was used to extrapolate room temperature  $E_g$  values to 0 K (blue curve) and temperature-dependent  $E_{PL}$  data fitted with same equation (red/black curves).

As temperature increases (up to  $\sim 80$  K) carriers are ionised and redistributed into other radiative and non-radiative defect states which quench PL emission (T>140 K). Activation energies of the defects involved in radiative recombination can be determined by analysis of the temperature-dependent PL spectra intensity using a two activation energy model proposed by (Luckert et al., 2011),

$$I(T) = \frac{I_0}{1 + C_1 exp\left(-\frac{E_{a1}}{kT}\right) + C_2 exp\left(-\frac{E_{a2}}{kT}\right)}$$
(6)

where  $I_0$  is the integrated intensity extrapolated to 0 K, k is the Boltzmann constant,  $C_1$  and  $C_2$  are the process rate parameters and  $E_{a1}$  and  $E_{a2}$  are the thermal activation energies for the involved defect states. Application of the model to the Arrhenius plots for both films (Fig. 4a) estimated activation

energies of  $E_{a1} = 40.1 \pm 6.4$  meV,  $E_{a2} = 5.9 \pm 0.5$  meV and  $E_{a1} = 38.3 \pm$ 280 3.8 meV,  $E_{a2} = 3.2 \pm 0.5$  meV for film HP and LP, respectively. The low 281 activation energies indicate the presence of shallow donor (3-6 meV) and 282 shallow acceptor (38-48 meV) states in both films. A kesterite defect study 283 by Chen et al. would suggest  $Sn_{Cu}$  and  $V_{Cu}$  are the corresponding donor 284 and acceptor states. They state the Sn 5p electron has a high orbital energy 285 which can be easily ionised, hence the  $Sn_{Cu}$  (0/+) defect is located just 286 below the conduction band minimum, which is in agreement with the very 287 shallow donor level observed in both types of absorber (Chen et al., 2013). 288 Also the ionised  $V_{Cu}$  defect is the predominant acceptor responsible for p-289 type conductivity in these CZTSSe absorbers producing a reasonably high 290 concentration of holes (in the region of  $10^{15}$ - $10^{16}$  cm<sup>-3</sup>). The presence of an 291 additional Sn-related acceptor defect  $(V_{Sn}, Zn_{Sn} \text{ or } Cu_{Sn})$  could explain the 292 anomalous red-shift in  $E_{PL}$  observed in film LP (see Fig. 3b). 293

The minority carrier dynamics of LP and HP CZTSSe thin films were 294 studied using time-resolved photoluminescence (TRPL) decays, see Fig. 4b. 295 Similar to PL, electron-hole (e-h) pairs are generated, in this case, by a short 296 pulse of incident photons and the time-dependence of the emitted light from 297 recombination of e-h pairs is monitored. The minority carrier lifetime  $\tau$  is 298 determined from the decay time of charge carrier recombination. In order 299 to evaluate  $\tau$ , the TRPL decay curves were fitted with a double exponential 300 function (Ohnesorge et al., 1998): 301

$$I_{PL}(t) = A_1 e^{-\left(\frac{t}{\tau_1}\right)} + A_2 e^{-\left(\frac{t}{\tau_2}\right)}$$
(7)

where  $I_{PL}(t)$  is the luminescence intensity at time t after the excitation pulse,  $A_1$  and  $A_2$  are are the PL intensities corresponding to the injection regimes and  $\tau_1$  and  $\tau_2$  are the fast and slow decay times. The initial fast decay  $\tau_1$  can be ascribed to high carrier injection immediately after the excitation pulse and the long tail  $\tau_2$  attributed to the minority carrier lifetime of the material (Gunawan et al., 2010; Repins et al., 2012).

Lifetimes of 0.20 and 0.44 ns were observed for films LP and HP, respec-308 The lower  $\tau$  value for film LP would suggest more non-radiative tively. 309 bulk/surface recombination centres are present in the absorber, which is 310 in agreement with the total defect density  $N_t$  determined from excitation-311 dependent PL measurements. Raadik et al. propose the shorter carrier life-312 time could be related to a higher degree of Cu/Zn disorder in the bulk of the 313 CZTSSe absorber (Raadik et al., 2017). Therefore different recombination 314 processes may be active in LP and HP films. The unexpectedly low carrier 315 lifetime at 6 K observed in this study could be due to high recombination 316 rates at the unpassivated absorber surface. 317

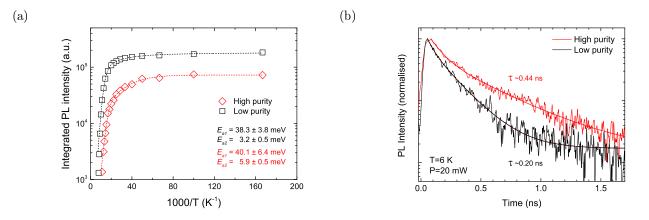


Figure 4: (a) Arrhenius plots of integrated PL intensities and (b) TRPL decays at 6 K for films LP and HP. The solid lines are results of fitting with bi-exponential function.

In order to probe deep defect levels in the CZTSSe devices beyond the

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scope of PL spectroscopy, deep level transient spectroscopy (DLTS) is em-319 ployed to quantify trap activation energies  $E_T$ , trap densities  $N_T$  and their 320 capture cross-sections  $\sigma_T$ . Supp. Fig. S2 show DLTS spectra for samples 321 LP and HP respectively. In this measurement setup positive  $\Delta C$  peaks are 322 indicative of hole trap levels with energy values measured with respect to the 323 valence band. Peak positions are analysed over three different transient pe-324 riod widths (19.2 ms, 192 ms, 480 ms) and for a range of correlator functions 325 (Weiss and Kassing, 1988), used to generate Arrhenius plots shown in Fig. 326 5a. Values for  $E_T$  and  $\sigma_T$  are extracted from fitting to the Arrhenius plot 327 whilst the trap density is determined from the magnitude of the capacitance 328 change relative to the reverse bias capacitance (see Table 2). 329

Table 2: List of electronic parameters of low and high purity CZTSSe samples.

		Trap level	Capture cross-	Trap density	Possible
		$E_T \ (\mathrm{meV})$	section $\sigma_T \ (\mathrm{cm}^2)$	$N_T~({ m cm}^{-3})$	$\operatorname{defect}$
Low purity	(1)	$86 \pm 7$	$(1.08 \pm 0.95) \ge 10^{-20}$	$(1.70 \pm 0.22) \ge 10^{14}$	$Cu_{Zn}$
	(2)	$167 \pm 10$	$(2.20 \pm 1.99) \ge 10^{-20}$	$(3.17\pm 0.32)\ge 10^{13}$	$Cu_{Sn}, Zn_{Sn}, V_{Sn}$
High purity	(1)	$87 \pm 6$	$(3.61 \pm 1.59) \ge 10^{-21}$	$(6.38 \pm 0.05) \ge 10^{14}$	$Cu_{Zn}$

For sample HP a single peak is clearly visible at an energy of  $E_T = 87 \pm$ 6 meV. The peak is very broad indicating the observed peak may in fact be a distribution of deep levels closely spaced in energy. Attempts were made to separate these levels via a Laplace analysis method (Dobaczewski et al., 2004) however such analysis did not give a definitive answer and was too dependent on the analysis conditions used i.e. number of overlapping transients

permissible. This sample also showed some evidence of a negative peak, in-337 dicative of an electron trap, at  $\sim 170$  K. On closer inspection, this was found 338 to predominantly occur for short period widths and is in fact measurement 339 artifact related to the recovery time of the capacitance signal. Sample LP 340 showed the same deep level present at  $E_T = 86 \pm 9$  meV, although the lower 341 purity sample displayed a slightly lower concentration of this trap level. A 342 second trap state not observable in the HP sample was also observed at an 343 energy of  $E_T = 168 \pm 12$  meV but an order of magnitude lower density, 3.17 344 x  $10^{13}~{\rm cm}^{-3},$  than for the shallower level . The natural assumption would 345 be that this secondary trap level has arisen as a result of some contaminant 346 within the process solution. It is assumed that the same trap level at 86-87 347 meV above the VBM in both types of absorber is antisite  $Cu_{Zn}$  (Chen et al., 348 2013). Although the concentration of this defect is higher in device HP than 349 LP, their potentially deleterious effect is mitigated by a lower carrier capture 350 cross-section ( $\sigma_{HP} \sim 4 \ge 10^{-21} \text{ cm}^2$  compared to  $\sigma_{LP} \sim 1 \ge 10^{-20} \text{ cm}^2$ ). Ab 351 initio calculations show that the additional defect level in film LP could be 352 Sn-related  $(V_{Sn}, Cu_{Sn} \text{ or } Zn_{Sn})$  with the likelihood of the defect being  $Zn_{Sn}$ , 353 considering the LP absorber is Cu-poor and Zn-rich (see Supp. Table S2). 354 The concentration of any Sn-related defects in film HP should be reduced as 355 the film is compositionally closer to the preferred Cu/(Zn+Sn) ratio of 0.8, 356 see Table 3) (Collord et al., 2015; Fairbrother et al., 2015). CZTSSe films 357 which have low Cu/(Zn+Sn) and high Zn/Sn ratios are predisposed due to 358 the presence of high populations of  $[Zn_{Sn} + 2Zn_{Cu}]$  charge compensated clus-359 ters which are one of the defect complexes responsible for non-stoichiometry 360 in this type of absorber material. 361

	Low purity	High purity
$\mathrm{Cu}/(\mathrm{Zn+Sn})$	0.84	0.82
Zn/Sn	1.07	1.08
$V_{OC}$ (V)	$0.31\ (0.31)$	$0.32 \ (0.32)$
$J_{SC}~({ m mA/cm^2})$	27.2(26.0)	25.6 (23.7)
<i>FF</i> (%)	59.9(57.9)	61.9(56.7)
Efficiency $\eta$ (%)	5.1 (4.7)	5.1(4.2)
Carrier lifetime $\tau$ (ns)	0.20	0.44
Carrier diffusion $L_d$ (nm) length	203	369

Table 3: Cu/(Zn+Sn) and Zn/Sn ratios of LP and HP thin films and solar cell parameters of subsequently fabricated best devices (with average values in brackets).

Further studies of the electronic properties of LP and HP solar cells were 362 performed using current density-voltage (J-V), capacitance-voltage (C-V)363 and external quantum efficiency (EQE) analysis. Fig. 5b shows the J-V plots 364 of the best performing LP and HP devices under 1-sun illumination with the 365 inset showing the average device open circuit voltage  $V_{OC}$  and short circuit 366 current density  $J_{SC}$ . It is noted that the average  $V_{OC}$  of HP cells was slightly 367 higher that of the LP cells, see inset of Fig. 5b. Conversely, LP devices 368 have a higher  $J_{SC}$  value (26.1 mA/cm<sup>2</sup>) in comparison to that seen in HP 369 devices  $(23.7 \text{ mA/cm}^2)$ . Similar observations were made by Yakushev et al. 370 for  $Cu_2ZnSnSe_4$  (CZTSe) devices with varying Cu/(Zn+Sn) and Zn/Sn ratios 371

and were attributed to the degree of Cu-Zn order/disorder in the crystal 372 lattice (Yakushev et al., 2017). EQE plots for LP and HP devices are shown in 373 Fig. 5c, where it can be seen that device LP has enhanced carrier extraction 374 in the wavelength range 600-1000 nm. This anomalous behaviour could be 375 explained by a lower concentration of free charge carriers and will be explored 376 in more detail later in this section. The minority carrier diffusion length 377  $L_d$  for both types of absorber was calculated from EQE measurements in 378 conjunction with optical absorption coefficient  $\alpha$  measurements (determined 379 from transmittance/reflectance data, see Supp. Fig. S3) using a method by 380 Coursel et al. (Coursel et al., 2016). It was shown that  $EQE^{-1}$  is a linear 381 function of  $\alpha^{-1}$  such that:-382

$$EQE(\lambda)^{-1} = \frac{1}{(1 - R(\lambda))} \left( 1 + \frac{\alpha(\lambda)^{-1}}{L_d} \right)$$
(8)

where  $R(\lambda)$  is the reflectance of the cell with an intercept on the  $\alpha^{-1}$  equal 383 to  $L_d$ , see Fig. 5d. Diffusion length values of 203 and 369 nm were de-384 termined for LP and HP devices respectively. The reported  $L_d$  values are 385 significantly lower than the value of 750 nm observed in the CZTSSe solar 386 cell with record efficiency of 12.6% (Wang et al., 2013). The longer diffusion 387 lengths of carriers in film HP strongly correlates with higher carrier lifetimes 388 measured in CZTSSe bulk. C-V measurements were performed to estimate 389 the depletion region width w and doping density  $N_A$  in the CZTSSe absorber 390 layers of devices LP and HP. The inset in Fig. 5d shows the plot of  $1/C^2$ 391 versus reverse bias voltage where  $\langle w \rangle = A \varepsilon_0 \varepsilon_r / C$  is determined from the 392 measured capacitance value at zero bias and  $N_A$  is derived from the slope 393  $d(1/C^2)/dV$ . Very short space charge regions widths of 63 and 61 nm were 394 observed for devices LP and HP respectively, with corresponding apparent 395

doping densities of 2.5 x10<sup>16</sup> and 3.3 x10<sup>16</sup> cm<sup>-3</sup>, in agreement with those observed by Qu et *al.* (Qu et al., 2018). A high doping density is usually associated with a short depletion width in CZTSSe solar cells as  $w \propto N_A$ (Ganchev et al., 2011; Haight et al., 2014).

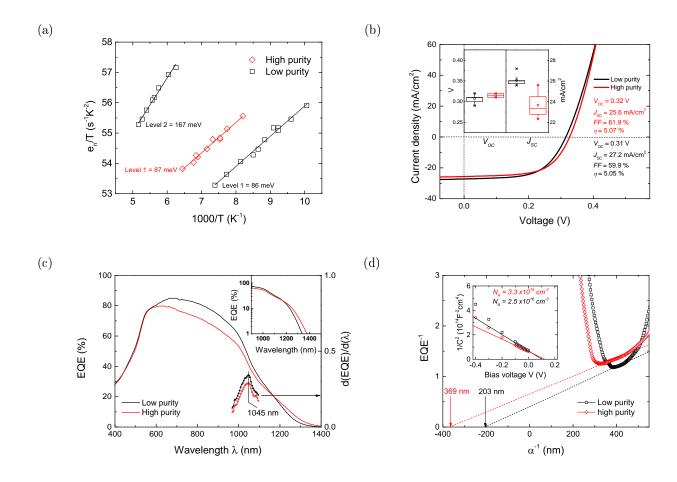


Figure 5: (a) Extracted Arrhenius plots from DLTS spectra, (b) J-V curves of best performing devices with inset showing box plots of average device  $V_{OC}$  and  $J_{SC}$  parameters, (c) EQE plots with inset showing extended absorption due to tail states in sample HP and (d) calculation of depletion region width  $w_0$  and apparent doping density  $N_A$  for samples LP and HP.

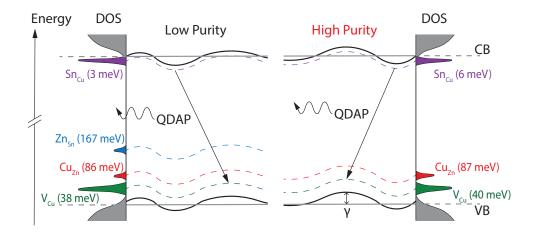


Figure 6: Schematic of electrostatic potential fluctuations in the band edges of LP and HP absorbers with associated defects and densities of state (DOS). The predominant radiative recombination process in both samples involves QDAP defects  $Sn_{Cu}$  and  $V_{Cu}$ .

Following the selenisation of the CZTS nanoparticle films, EDS compo-400 sitional analysis of the LP and HP CZTSSe absorbers revealed a greater Cu 401 and Sn loss was observed in LP films, whereas HP films suffered a larger Zn 402 loss. Subsequent Cu/Zn+Sn and Zn/Sn ratios of 0.84, 0.82 and 1.07, 1.08 for 403 respective LP and HP absorbers demonstrate HP films are closer to the pre-404 ferred CZTSSe composition of Cu/Zn+Sn = 0.8 and Zn/Sn = 1.2 (Collord 405 et al., 2015; Fairbrother et al., 2015). With regard to the elemental losses, 406 it would be reasonable to assume that there could be higher concentrations 407 of Cu- and Sn-related defects in LP films. Recent studies have shown the 408 degree of Cu-Zn order/disorder has a significant impact on the bandgap and 409 crystallinity of CZTSSe absorbers and may be the main cause of bandgap 410 and electrostatic potential fluctuations within the material, as shown in Fig. 411 6 (Rey et al., 2014; Scragg et al., 2015; Rey et al., 2018; Schorr, 2011). 412

Considering the small decrease in Cu/Zn+Sn and small increase in Zn/Sn 413 ratios in LP films compared to HP films, higher concentrations of A-type de-414 fects ( $[V_{Cu}+Zn_{Cu}]$  neutral complexes) may be present in HP films together 415 with higher concentrations of B-type defects  $([Zn_{Sn}+2Zn_{Cu}])$  neutral com-416 plexes) in LP films (Lafond et al., 2012; Gurieva et al., 2015). Although the 417 presence of  $Zn_{Cu}$  donor defect was not directly observed in either film, their 418 presence can be inferred as high populations of benign  $[V_{Cu}+Zn_{Cu}]$  clusters 419 are expected in non-stoichiometric CZTSSe absorbers (Chen et al., 2013). 420

Perhaps the most interesting observation in this study is the presence of 421 an additional acceptor defect in the LP device located  $\sim 167$  meV above the 422 VBM, speculatively attributed to antisite  $Zn_{Sn}$ . Chen et al. have shown the 423  $Zn_{Sn}$  defect has negligible impact on p-type conductivity but could contribute 424 to luminescence and act as recombination centres (Chen et al., 2013). The 425 presence of radiative  $Zn_{Sn}$  defects could explain the increased QDAP den-426 sity and PL intensity seen in film LP compared to film HP. The significant 427 red-shift of the PL peak in film LP at temperatures greater than 80 K could 428 be attributed to radiative recombination involving the  $Zn_{Sn}$  acceptor. The 429 lower QDAP concentration in film HP could also account for quenching of 430 the PL signal at a lower temperature ( $\sim 100$  K) in this film compared to film 431 LP ( $\sim 140$  K). It is also possible the anomalous temperture-dependent PL be-432 haviour in film LP could involve  $Cu_{Zn}$ - $Sn_{Cu}$  QDAP due to the thermal escape 433 of holes from the shallower  $V_{Cu}$  acceptor state to the slightly deeper  $Cu_{Zn}$ 434 state, see Fig. 6. According to Chen et al. a number of self-compensated 435 defect clusters can be formed in quaternary kesterites. They show the overall 436 formation energy of these defect clusters is significantly decreased relative to 437

the sum of the individual defects. As the clusters are charge-nuetral and 438 bound by strong Coulomb attraction, their presence should not be detected 439 directly by PL analysis. However, some defect clusters can produce a sig-440 nificant shift in the valence and conduction band edges, effectively reducing 44: the bandgap of the absorber. Large populations of clusters  $[2V_{Cu}+Sn_{Zn}]$ 442 or  $[2Cu_{Zn}+Sn_{Zn}]$  could exist in both types of absorber given the observed 443 large red-shift of  $E_{PL}$  in comparison to the bandgap of the absorbers at 6 444 K ( $\sim 300 \text{ meV}$ ) which cannot be solely attributed to QDAP defects. The 445 presence of neutral defect clusters would reduce the concentration of shallow 446 acceptor defects which contribute to the overall p-type conductivity of the 447 absorber material. From capacitance-voltage measurements, film LP has a 448 slightly lower apparent doping density  $N_A$  in comparison to that of film HP 449 indicating the degree of compensation is greater in this film. The measured 450 concentration of  $V_{Cu}$  antisite defects (which are mainly responsible for the 451 p-type conductivity of Cu-poor/Zn- rich CZTSSe films) is around  $10^{19}~{\rm cm^{-3}}$ 452 in both types of absorber whereas both films show  $N_A \sim 10^{16} {\rm ~cm^{-3}}$  confirming 453 the material is highly compensated. 454

All optical and electronic parameters indicate films prepared from high 455 purity precursor chemicals are of a higher quality than those fabricated from 456 lower grade materials. However, this improvement in film quality does not 457 translate to an increase in completed device efficiency as both types of device 458 have similar best  $\eta$  values of ~5.1 %. A previous study of CZTSSe solar 459 cells fabricated using the same nanoparticle inks-based method has revealed 460 the devices typically have a contact barrier height of  $\sim 40$  meV between the 461 CZTSSe absorber and Mo back contact (Campbell et al., 2018). This low 462

<sup>463</sup> barrier height would suggest the back contact of both types of device is not
<sup>464</sup> a factor constraining cell performance and, therefore, attention should be
<sup>465</sup> focused on improving buffer/absorber junction and other components in the
<sup>466</sup> device architecture.

One important parameter which could be utilised to indicate the quality of 467 CZTSSe absorbers is the mean depth of potential fluctuations  $\gamma$ . Considering 468 a high concentration of  $Zn_{Sn}$  defects present in absorber LP (which could 469 contribute to an elevated QDAP density), we see a direct correlation of such 470 QDAP density with  $\gamma$ , suggesting  $\gamma$  is a useful indicator of the presence of 471 specific defects which generate band tails in the material. Yakushev et al. 472 observe a similar reduction in  $\gamma$  values when the studied CZTS films became 473 Cu-poorer and Zn-richer, attributing this change to improved ordering of 474 Cu-Zn atoms within the crystal lattice of the thin film (Yakushev et al., 475 2017). A similar reduction in  $\gamma$  values are noted in this study. Bourdais et 476 al. investigated the effect of Cu-Zn disorder on the  $V_{OC}$  deficit in kesterite 477 solar cells also noted that ordering of the CZTSSe absorbers by slow cooling 478 improved the  $V_{OC}$  deficit in their devices by 40 meV. Therefore, a reduction 479 in  $\gamma$  value could indicate a reduction in Cu-Zn disorder with an associated 480 increase in  $V_{OC}$  (Bourdais et al., 2016). Fig. 6 highlights the main results of 481 the combined optical and electronic studies, principally: 482

(i) Fluctuations in the band edges of both types of absorber, predominantly caused by variations in electrostatic potential, with an average depth of  $\gamma$  equal to 59 and 49 meV in films LP and HP respectively

(ii) Same shallow donor  $(Sn_{Cu})$  and shallow acceptor  $(V_{Cu}, Cu_{Zn})$  defects present in the bulk of both types of CZTSSe absorber (iii) Additional deep Sn-related acceptor defect (probably  $Zn_{Sn}$ ) located 167 meV above the VBM present in sample LP

(iv) Radiative recombination primarily involving QDAP donor  $Sn_{Cu}$  and acceptor  $V_{Cu}$  defects located 3-6 meV below CBM and 38-40 meV above VBM

(v)  $V_{Cu}$  chiefly responsible for *p*-type conductivity in both types of thin film.

#### 495 4. Conclusion

In depth optical spectroscopy and electronic studies of CZTSSe absorber 496 layers fabricated from low and high purity precursor chemicals are presented. 497 The high purity chemical recipe produced Cu-poorer and Zn-richer CZTS 498 nanoparticles. Following selenisation of the stacked spin-coated CZTS films, 499 a Zn loss was observed in the HP absorber while a Sn loss was seen in the LP 500 absorber. Comprehensive temperature and excitation dependent analysis of 501 PL spectra together with DLTS enabled identification of shallow and deep 502 defects and QDAP recombination as the predominant recombination mecha-503 nism in both types of thin film. The loss of Sn could account for the presence 504 of additional Sn-related defects identified in the LP film. Notwithstanding 505 the improvement in quality of film HP, detailed analysis of the electronic 506 properties of LP and HP solar cells revealed similar performance in both 507 types of device. This implies an underlying issue other than defects in the 508 absorber bulk inhibits device performance which warrants further investiga-509 tion. 510

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## <sup>694</sup> Appendix A. Supplementary material

Table S1: Chemical composition and concentration details of low/high purity CZTS nanoparticle recipe.

	Low puri	ty recipe	High	High purity recipe	
Copper(II) acetylacetonate $(Cu(acac)_2)$	97%	Sigma Aldrich	99.99%	Sigma Aldrich	
$Zinc \\ acetylacetonate \\ (Zn(acac)_2)$	not provided by supplier	Alfa Asear	99.995%	Sigma Aldrich	
Tin(IV) bis(acetylacetonate) dichloride (Sn(acac)_2Cl_2)	95%	Alfa Asear	98%	Sigma Aldrich	
$Elemental \ sulphur \ (S)$	99.98%	Sigma Aldrich	99.98%	Sigma Aldrich	
Oleylamine (OLA)	technical grade 70%	Sigma Aldrich	98%	Sigma Aldrich	

		Cu	Zn	Sn	S	Se	$\frac{Cu}{Zn+Sn}$	Zn/Sn	$\frac{Se}{Se+S}$
Solvent		(at%)	(at%)	(at%)	(at%)	(at%)			
HP	Precursor	23.93	16.88	14.24	45.95		0.74	1.19	
	Selenised	20.61	13.13	12.12	5.57	48.53	0.82	1.08	0.90
LP	Precursor	23.17	14.3	14.16	48.04		0.80	1.03	
	Selenised	19.09	11.80	11.02	5.02	53.06	0.84	1.07	0.91

Table S2: The composition of precursor and selenised films on bare SLG

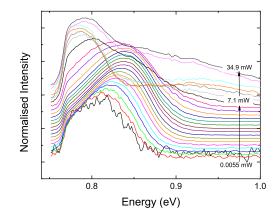


Figure S1: Emergence of high energy 'shoulder' in PL spectra of film HP indicating saturation of QDAP defects and an increasing contribution to PL intensity from band-related radiative recombination.

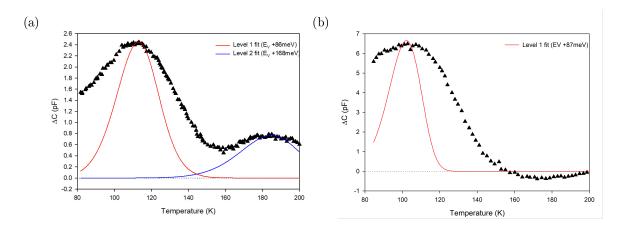


Figure S2: DLTS spectra recorded for (a) sample LP and (b) sample HP.

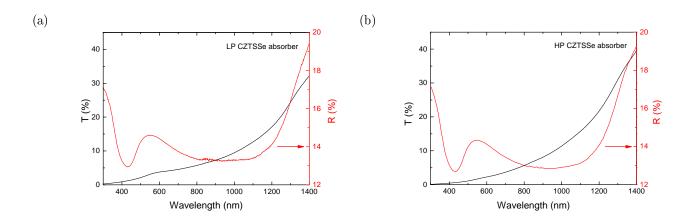


Figure S3: Transmission and absorption data of CZTSSe absorbers deposited on bare soda lime glass of (a) film LP and (b) film HP.